

Impact of GNSS Orbit Modelling on Reference Frame Parameters

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Overview

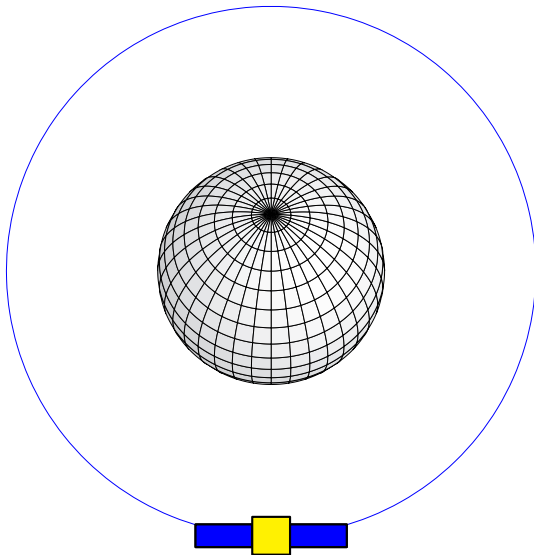
Solar Radiation Pressure for GNSS Satellites

Impact on the GNSS Satellite orbits

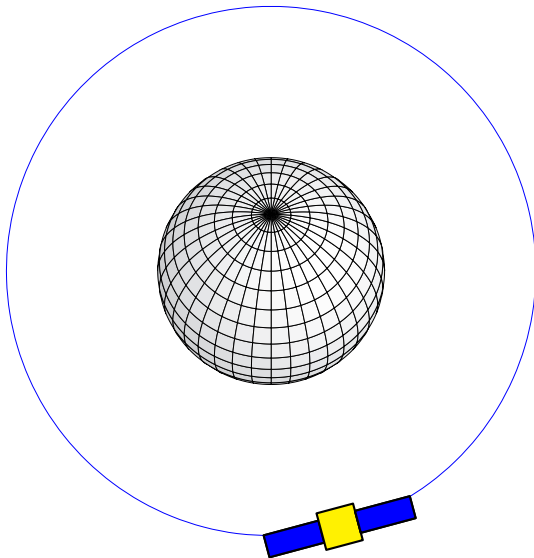
Impact on the Reference Frame Parameters

Long-Arc Solutions

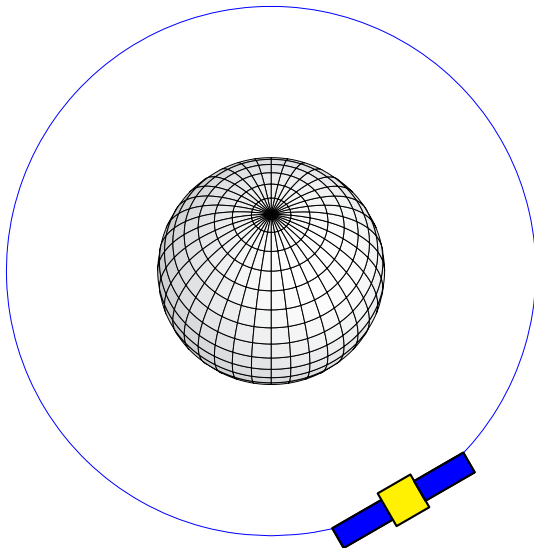
Observing the satellite from the Sun



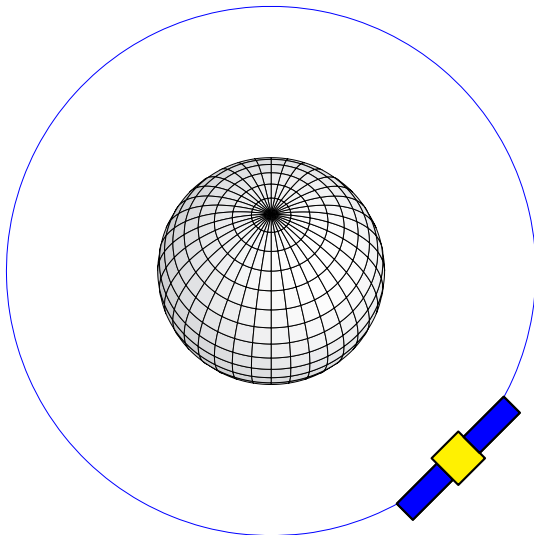
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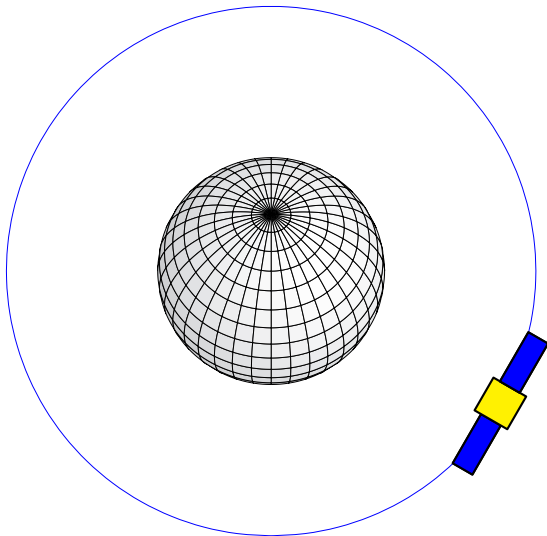
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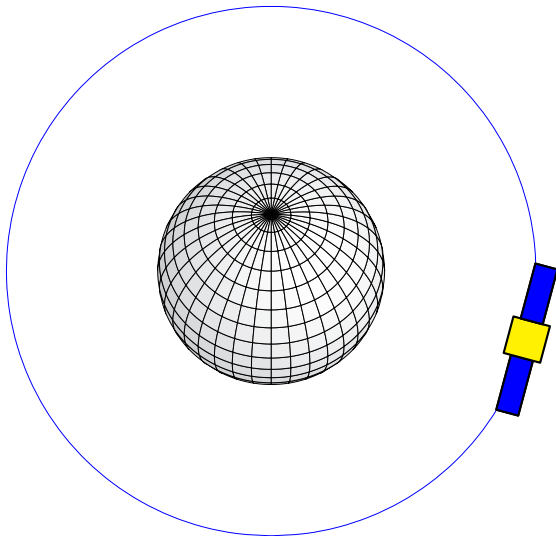
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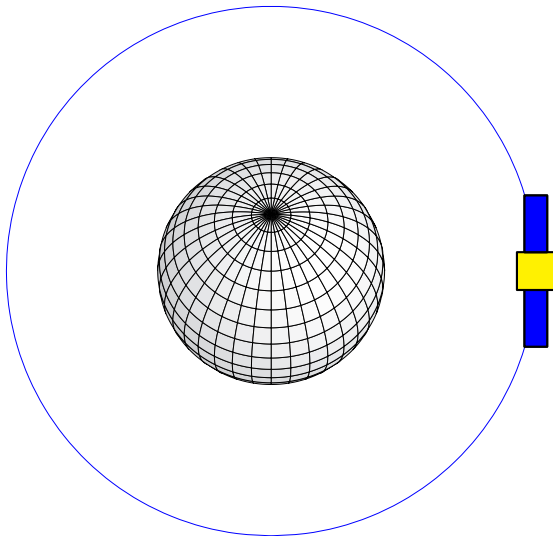
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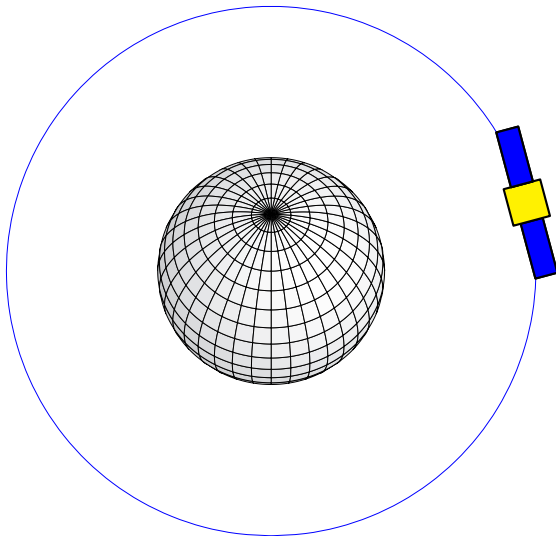
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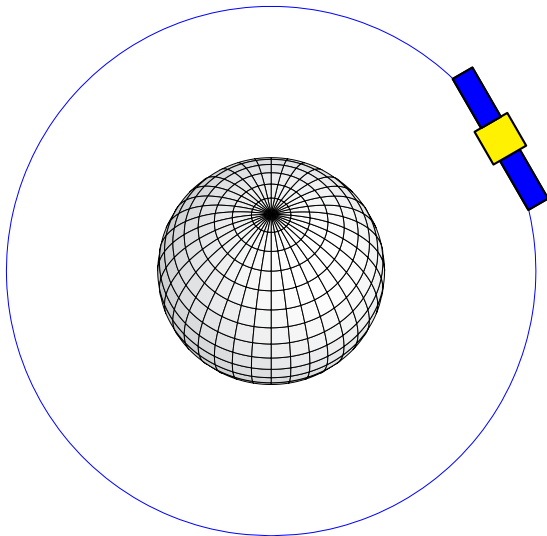
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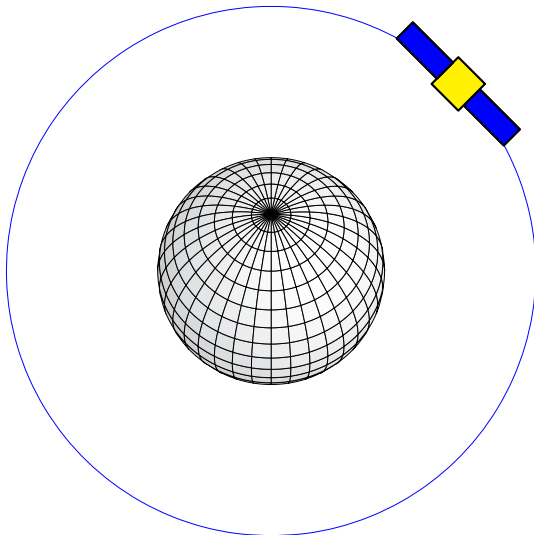
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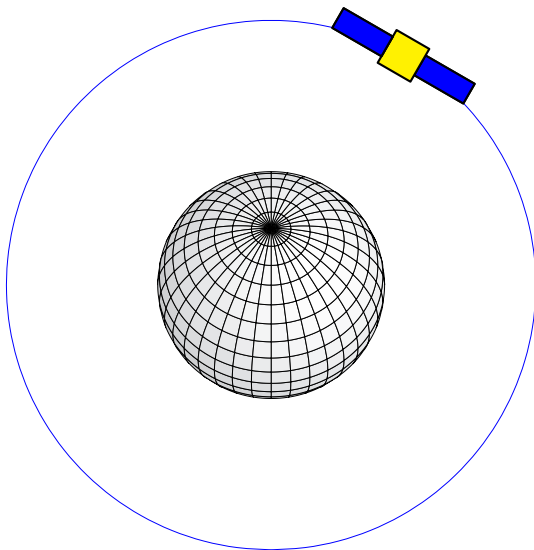
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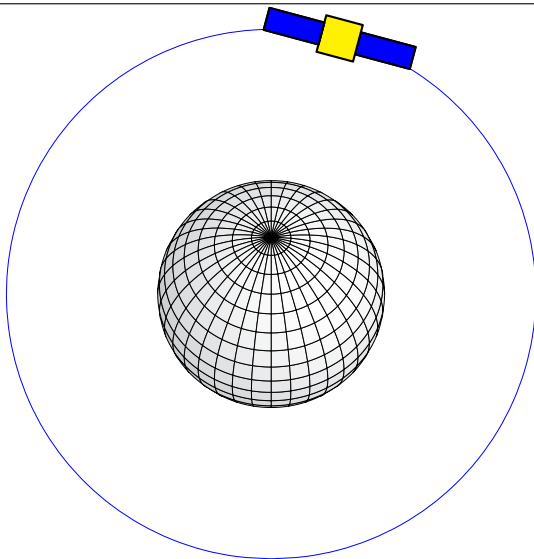
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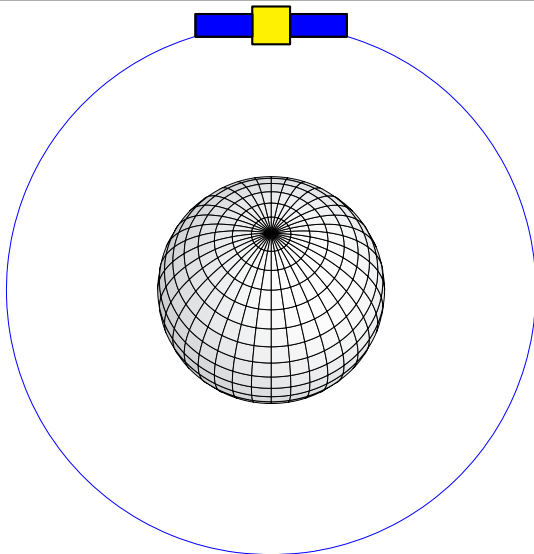
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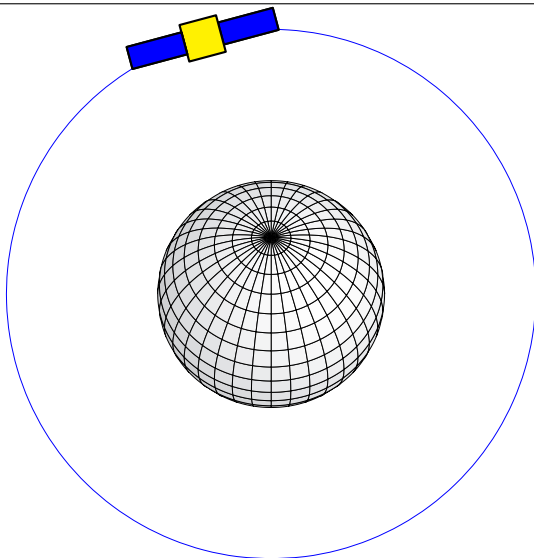
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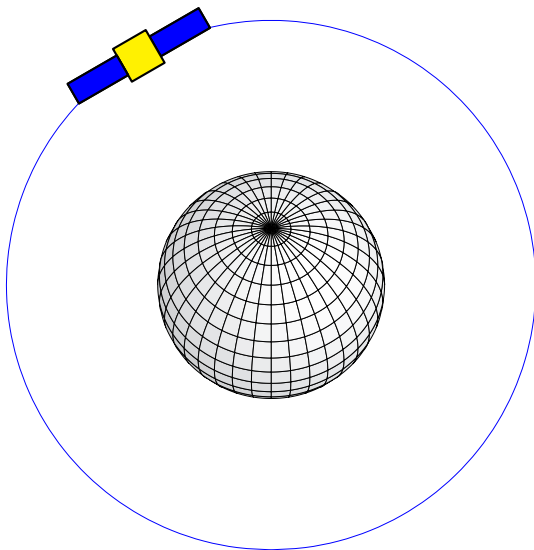
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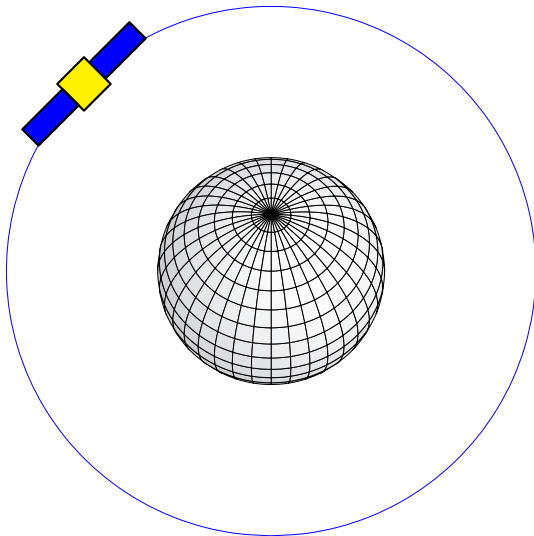
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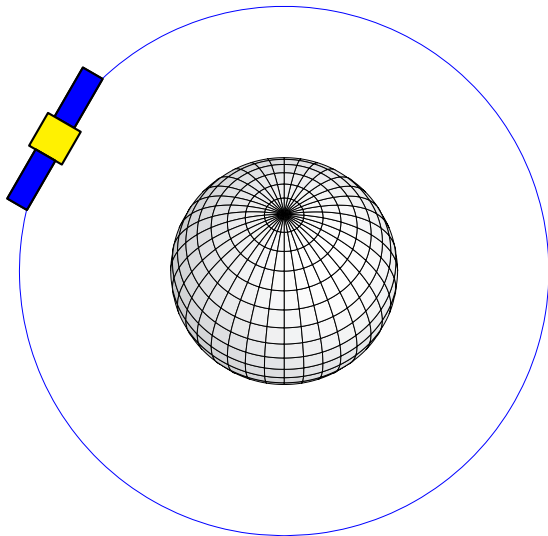
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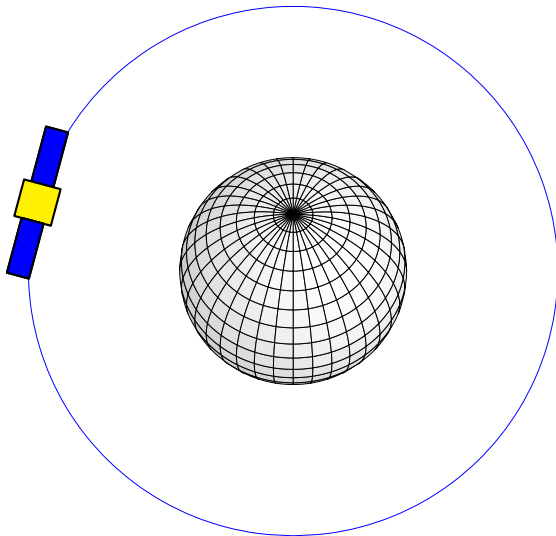
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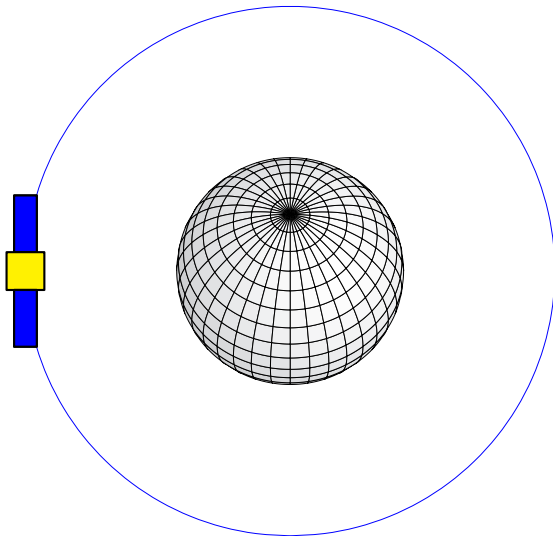
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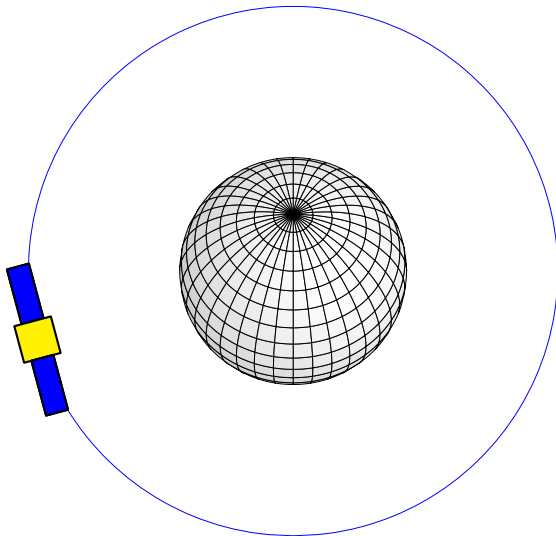
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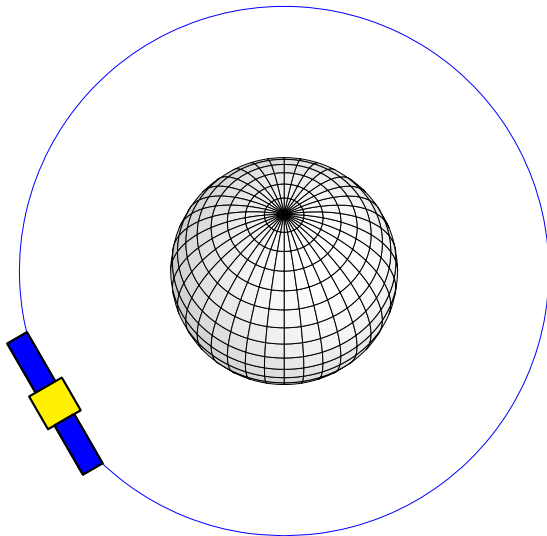
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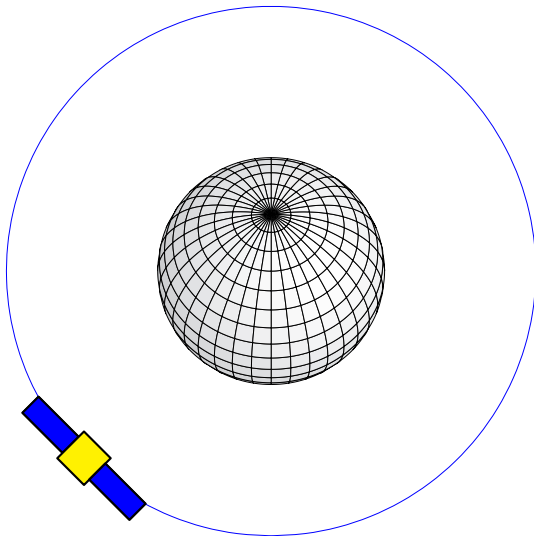
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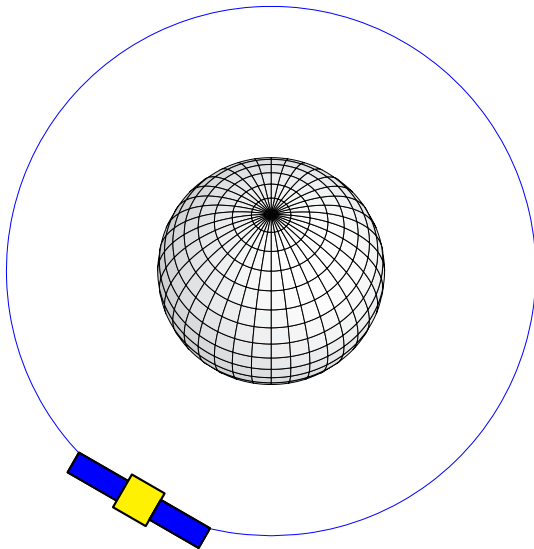
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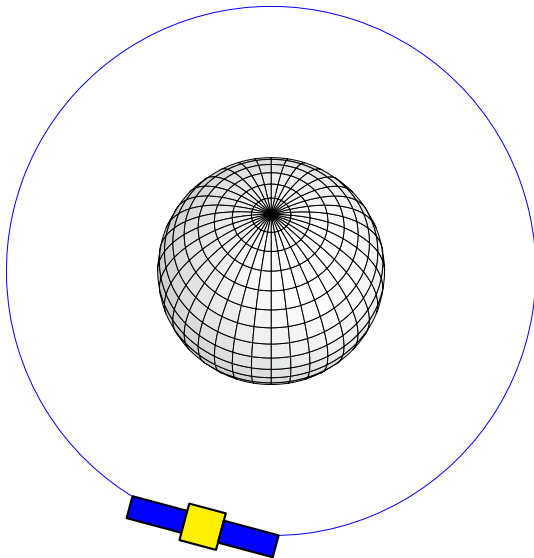
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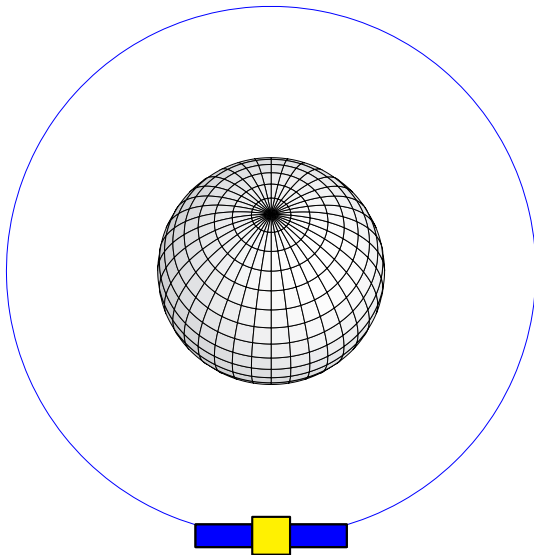
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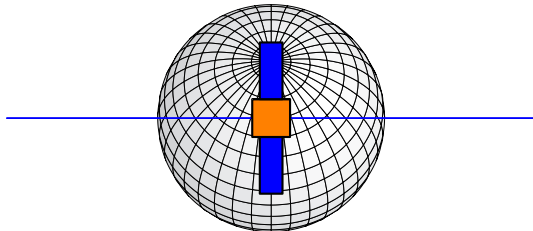
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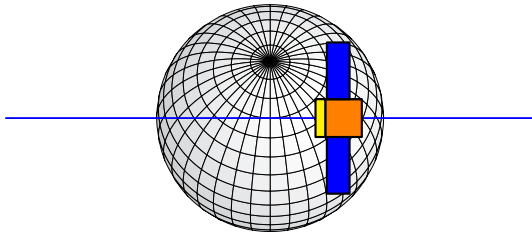
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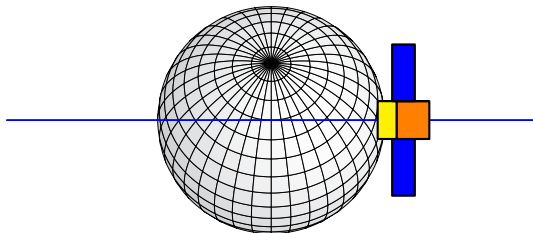
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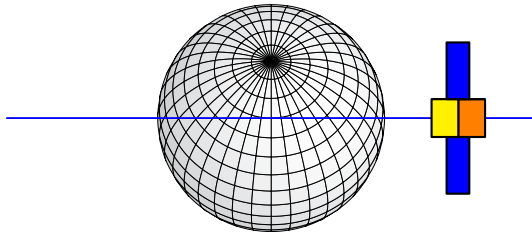
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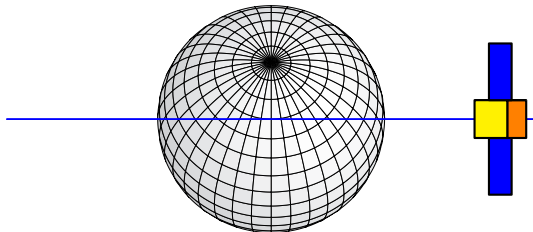
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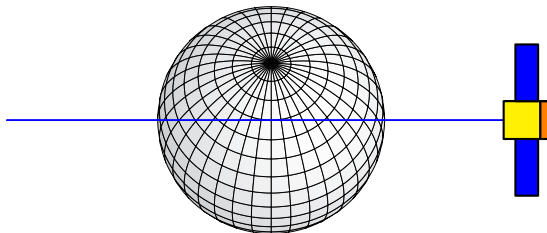
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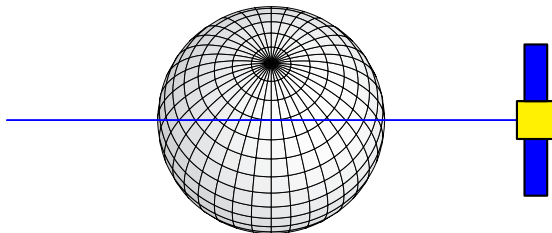
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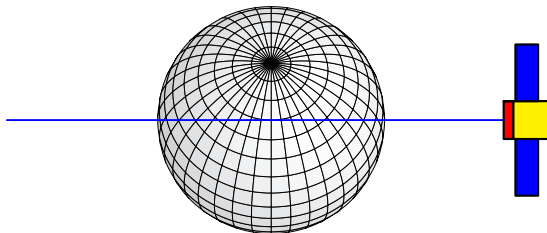
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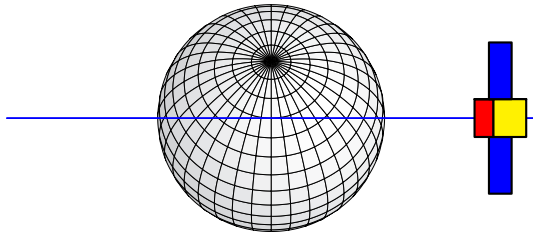
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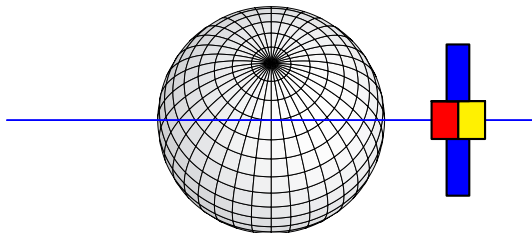
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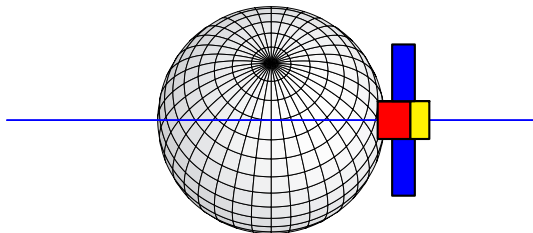
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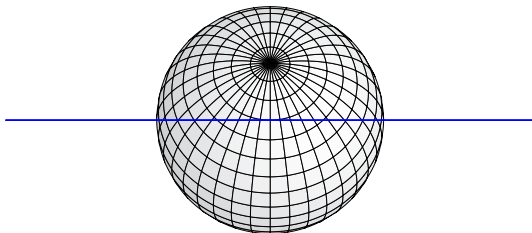
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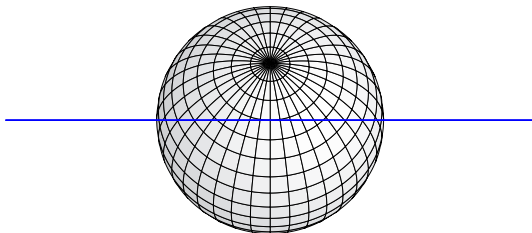
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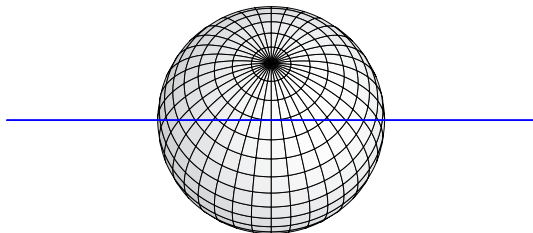
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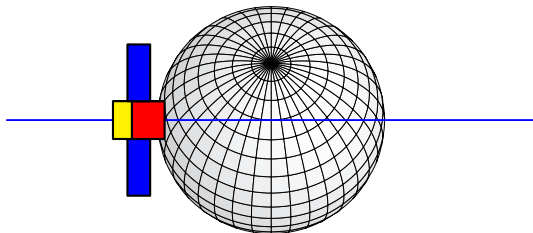
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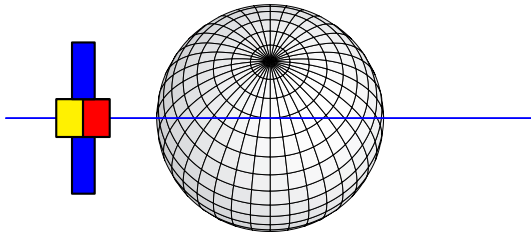
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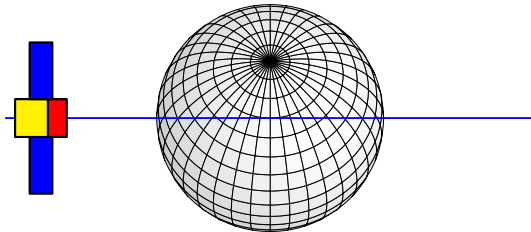
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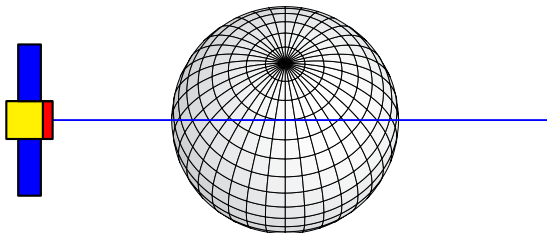
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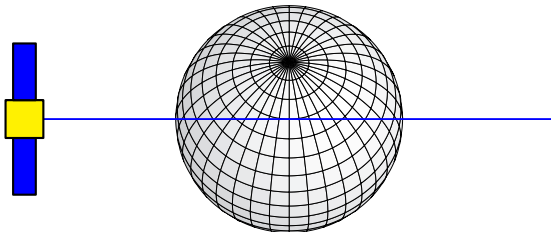
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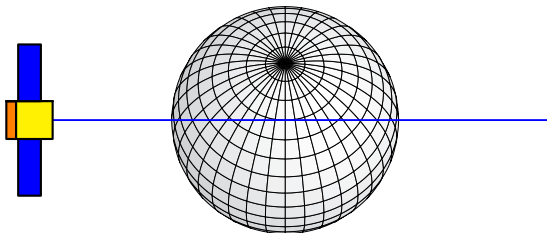
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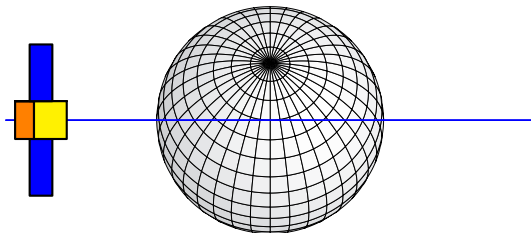
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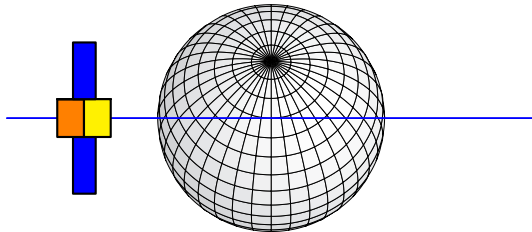
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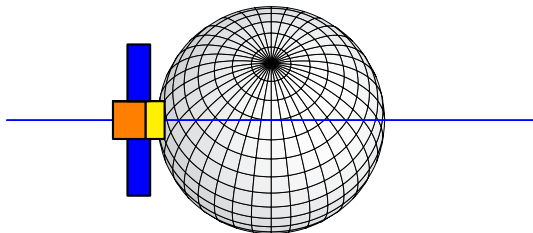
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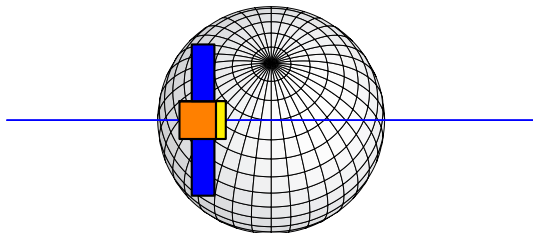
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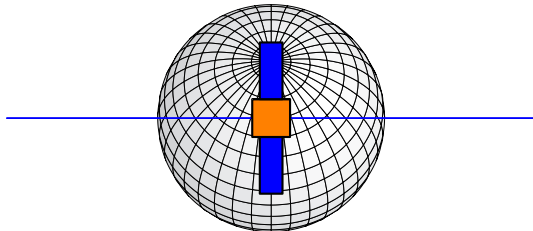
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Conclusions

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- If the Sun is perpendicular to the orbital plane no periodic solar radiation pressure perturbations are expected.
- If the Sun is located in the orbital plane a once-per-revolution signal is expected in the X -direction and a twice-per-revolution signal in the D -direction.

Observing the satellite from the Sun

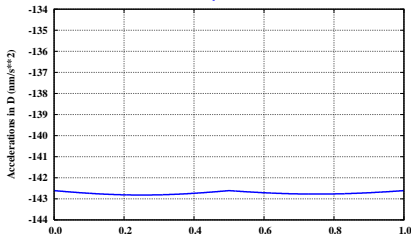
Conclusions

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- If the Sun is perpendicular to the orbital plane no periodic solar radiation pressure perturbations are expected.
- If the Sun is located in the orbital plane a once-per-revolution signal is expected in the X -direction and a twice-per-revolution signal in the D -direction.
- These periodic signals are the more pronounced the more the satellite body deviates from a sphere
(less for a cube – GPS – than a cylinder – GLONASS)

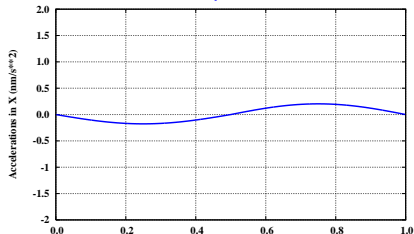
Solar radiation pressure from models

Accelerations derived for GLONASS satellites from a boxwing model¹

D component



X component



Computed for

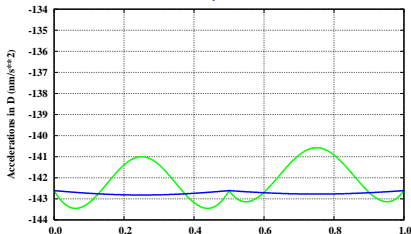
$$\beta = 88^\circ$$

¹as proposed by Carlos Rodríguez-Solano for the IGS based on Ziebart (2001)

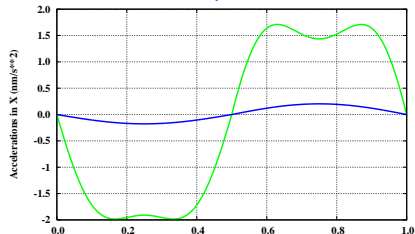
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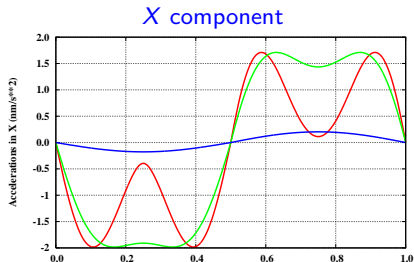
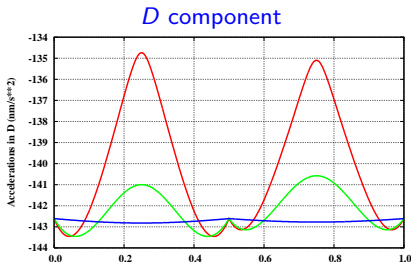
$\beta = 45^\circ$

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Computed for

$\beta = 10^\circ$

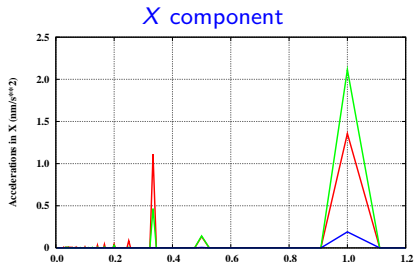
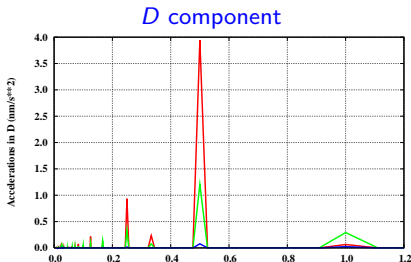
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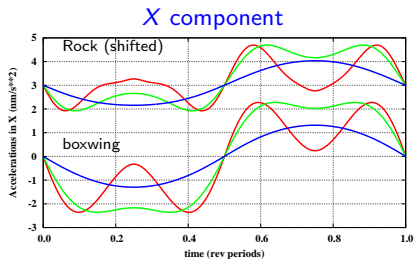
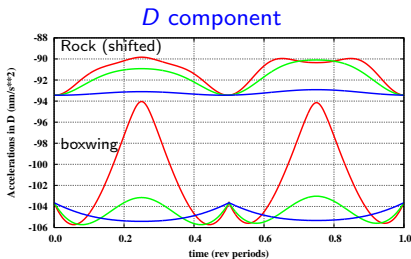
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Solar radiation pressure from models

Accelerations derived for GPS (Block IIA) satellites from a boxwing² and Rock-S³ model



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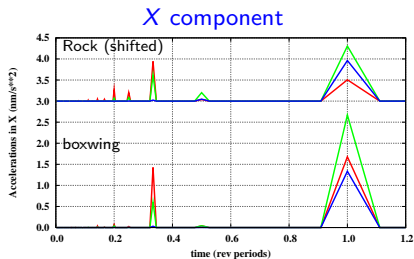
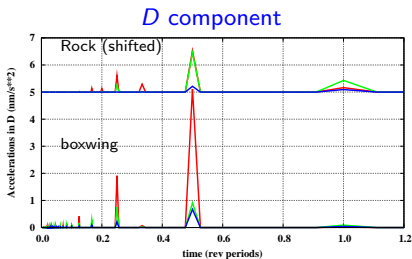
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Conclusions

- A Sun-fixed argument for the periodic terms is necessary to obtain interpretable series of these parameters:

$$u' = u_{sat} - u_{Sun}$$

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- A Sun-fixed argument for the periodic terms is necessary to obtain interpretable series of these parameters:

$$u' = u_{sat} - u_{Sun}$$

- Solar radiation pressure for satellites flying according to the previously mentioned models can be represented by:

$$D = D_0 + D_2 \cos(2u') + D_4 \cos(4u') + \dots$$

$$Y = (Y_0)$$

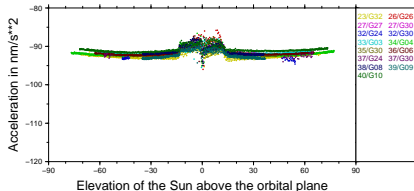
$$X = X_1 \cos(1u') + X_3 \cos(3u') + \dots$$

$Y_0 \neq 0$ if the satellite is flying “missaligned” with a Y -bias (e.g., GPS, except for Block IIF).

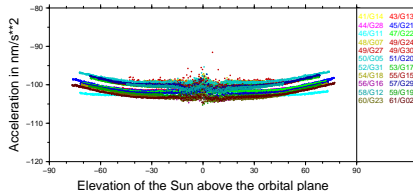
Estimated solar radiation pressure

Component: D_0

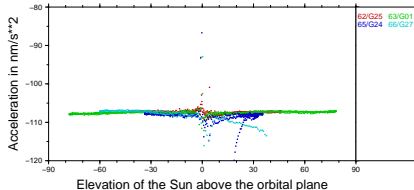
GPS Block IIA



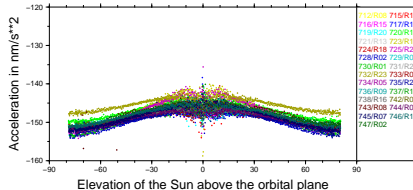
GPS Block IIR



GPS Block IIF



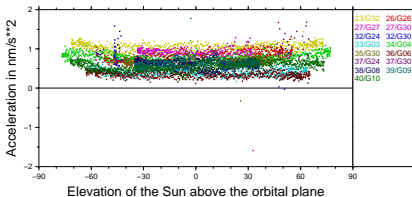
GLONASS-M



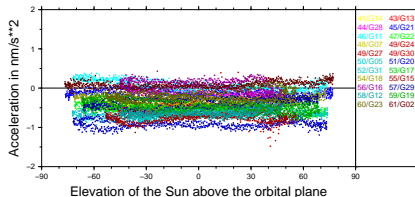
Estimated solar radiation pressure

Component: Y_0 (small scale)

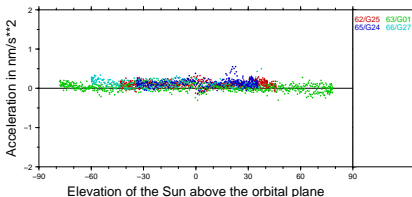
GPS Block IIA



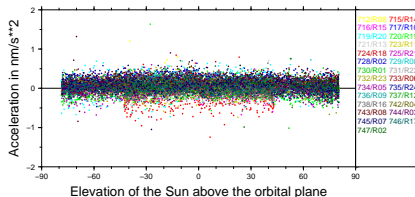
GPS Block IIR



GPS Block IIF



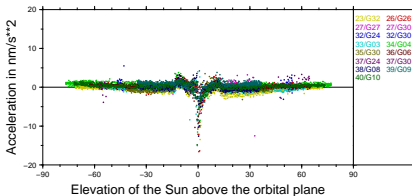
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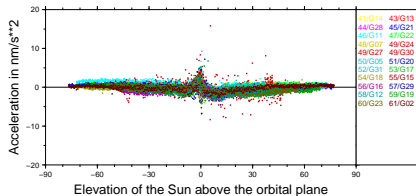
Estimated solar radiation pressure

Component: $X_1 \cdot \cos(1u')$

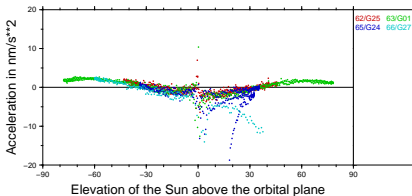
GPS Block IIA



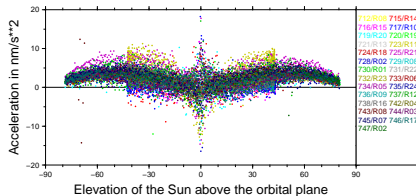
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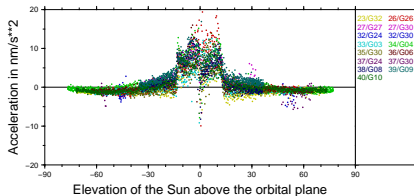
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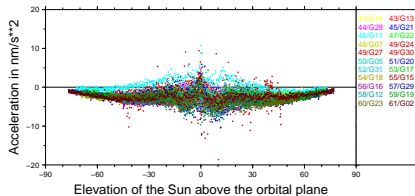
Estimated solar radiation pressure

Component: $D_2 \cdot \cos(2u')$

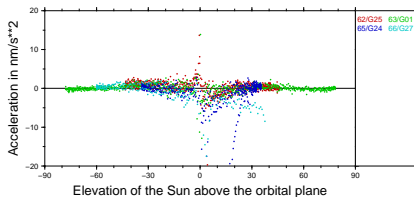
GPS Block IIA



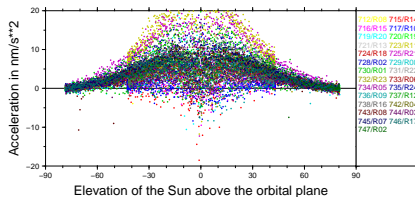
GPS Block IIR



GPS Block IIF



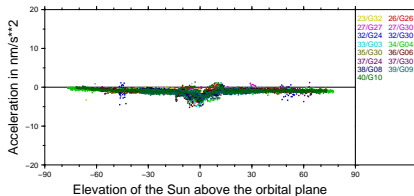
GLONASS-M



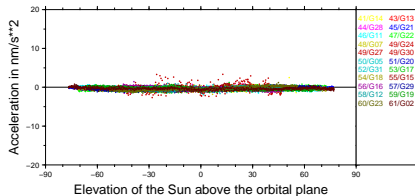
Estimated solar radiation pressure

Component: $X_1 \cdot \sin(1u')$

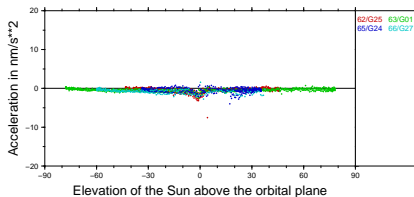
GPS Block IIA



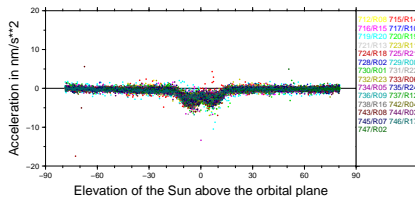
GPS Block IIR



GPS Block IIF



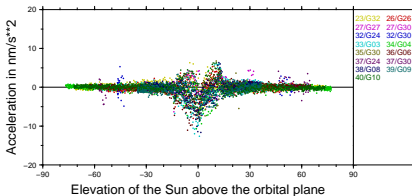
GLONASS-M



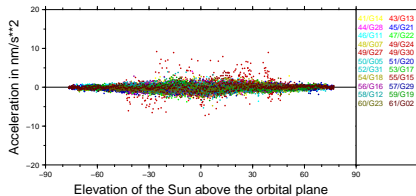
Estimated solar radiation pressure

Component: $D_2 \cdot \sin(2u')$

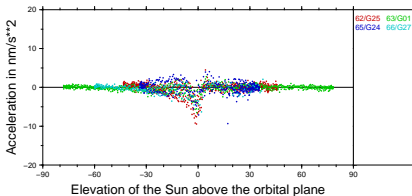
GPS Block IIA



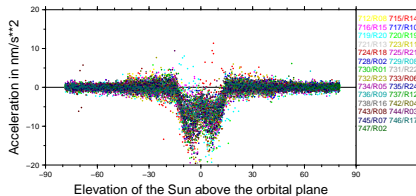
GPS Block IIR



GPS Block IIF

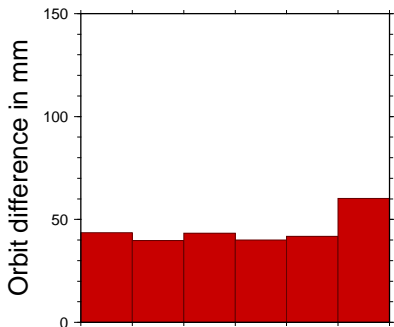


GLONASS-M



Impact on the GNSS Satellite orbits

Orbit overlaps from one-day solutions (mean over all components)

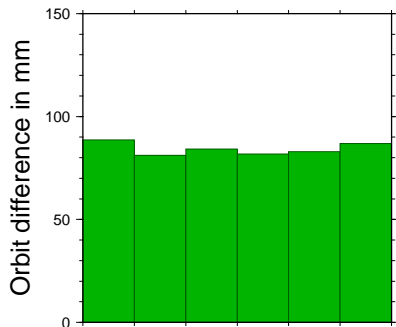


Solutions (GPS)

Components of the orbit model:

X

D



Solutions (GLONASS)

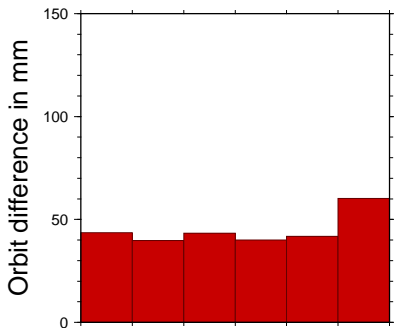
Components of the orbit model:

X

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Impact on the GNSS Satellite orbits

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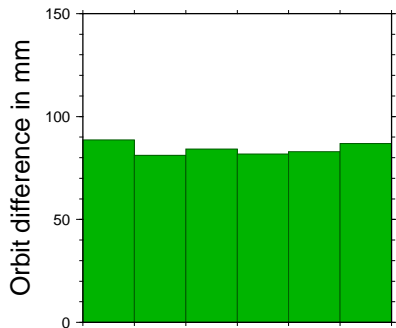


Solutions (GPS)

Components of the orbit model:

X $1u'$

D $-$



Solutions (GLONASS)

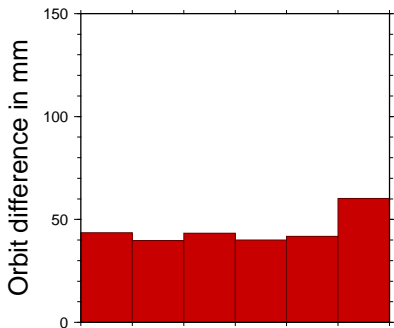
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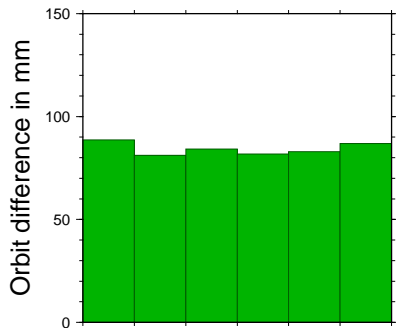


Solutions (GPS)

Components of the orbit model:

X $1u'$ $1u'$

D $-$ $2u'$



Solutions (GLONASS)

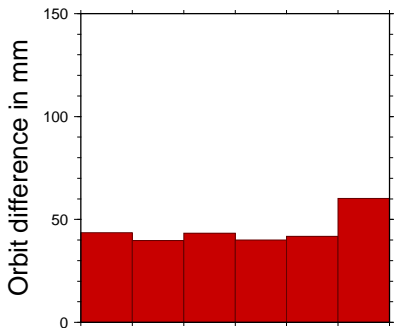
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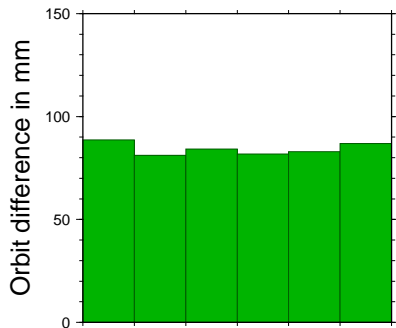


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Components of the orbit model:

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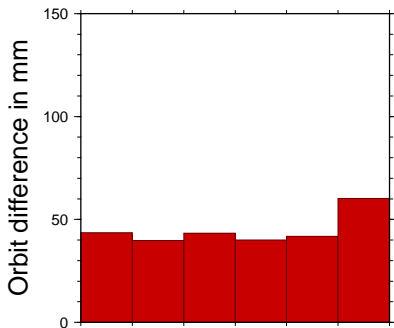
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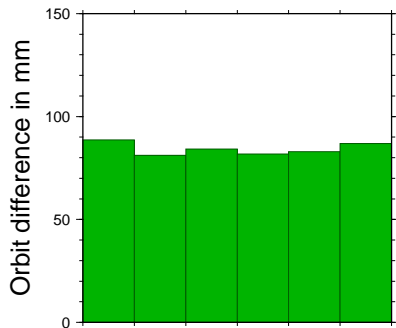


Solutions (GPS)

Components of the orbit model:

$X \quad 1u' \quad 1u' \quad - \quad 1u'$

$D \quad - \quad 2u' \quad 2u' \quad 2u'$
 $4u'$



Solutions (GLONASS)

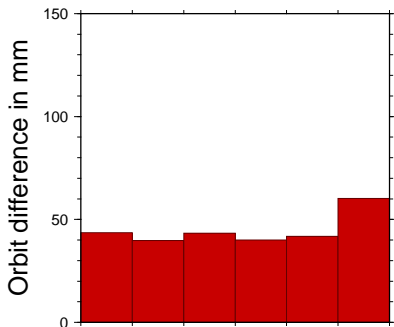
Components of the orbit model:

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 $4u'$

Impact on the GNSS Satellite orbits

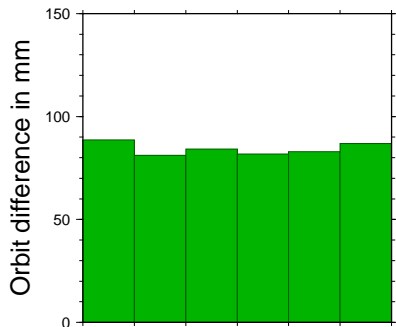
Orbit overlaps from one-day solutions (mean over all components)



Solutions (GPS)

Components of the orbit model:

X	$1u'$	$1u'$	$-$	$1u'$	$1u'$
				$3u'$	
D	$-$	$2u'$	$2u'$	$2u'$	$2u'$
				$4u'$	



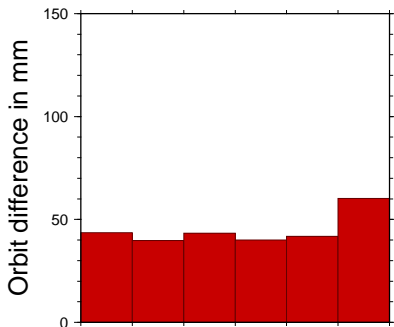
Solutions (GLONASS)

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X	$1u'$	$1u'$	$-$	$1u'$	$1u'$
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Impact on the GNSS Satellite orbits

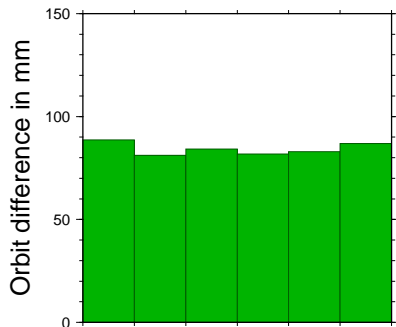
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Solutions (GPS)

Components of the orbit model:

X	1u'	1u'	-	1u'	1u'	1u'
				3u'	3u'	
D	-	2u'	2u'	2u'	2u'	2u'
				4u'		(cos)



Solutions (GLONASS)

Components of the orbit model:

X	1u'	1u'	-	1u'	1u'	1u'
				3u'	3u'	
D	-	2u'	2u'	2u'	2u'	2u'
				4u'		(cos)

Conclusions

- The new definition of the angular argument ($u' = u_{sat} - u_{Sun}$ instead of u_{sat}) allows it to better interpret of estimated parameter series, e.g., w.r.t. the elevation of the Sun above the orbital plane.

Conclusions

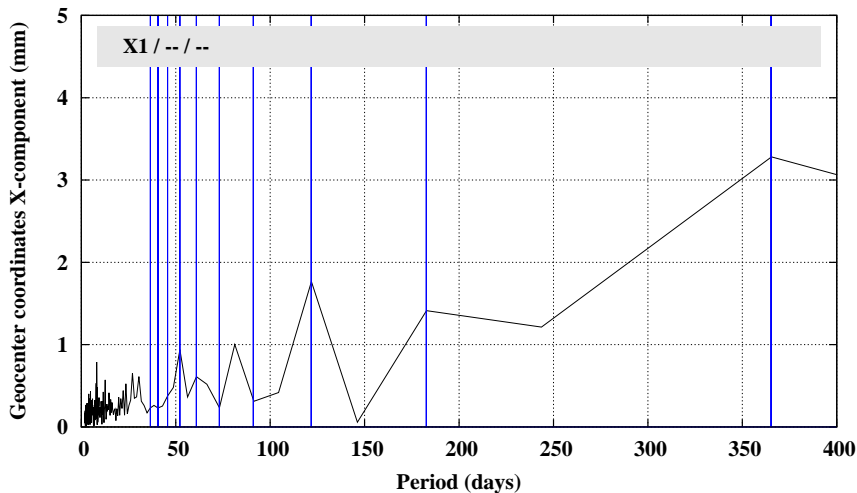
- The new definition of the angular argument ($u' = u_{sat} - u_{Sun}$ instead of u_{sat}) allows it to better interpret of estimated parameter series, e.g., w.r.t. the elevation of the Sun above the orbital plane.
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Conclusions

- The new definition of the angular argument ($u' = u_{sat} - u_{Sun}$ instead of u_{sat}) allows it to better interpret of estimated parameter series, e.g., w.r.t. the elevation of the Sun above the orbital plane.
- Adding twice-per-revolution terms in D -component improves the orbit solution.
- Even if the sin-terms are not necessary according to theory they are needed for representing real satellite trajectories.

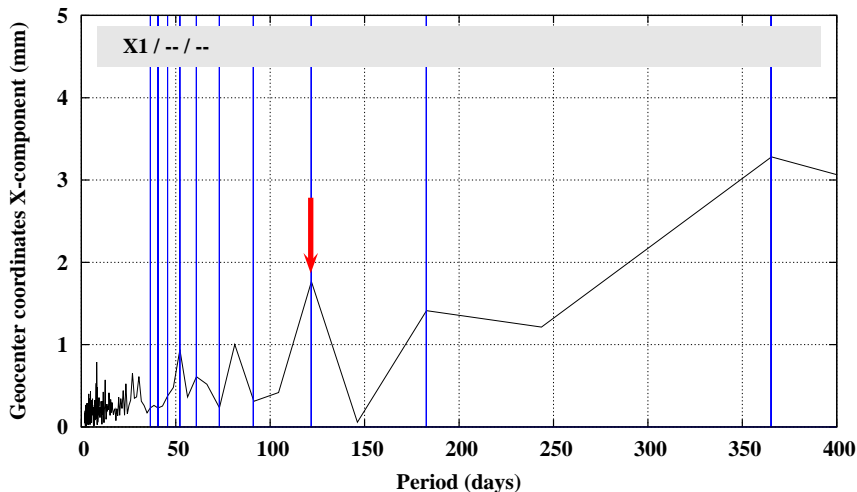
Impact on the Geocenter Estimates

Spectra from geocenter estimates: X component



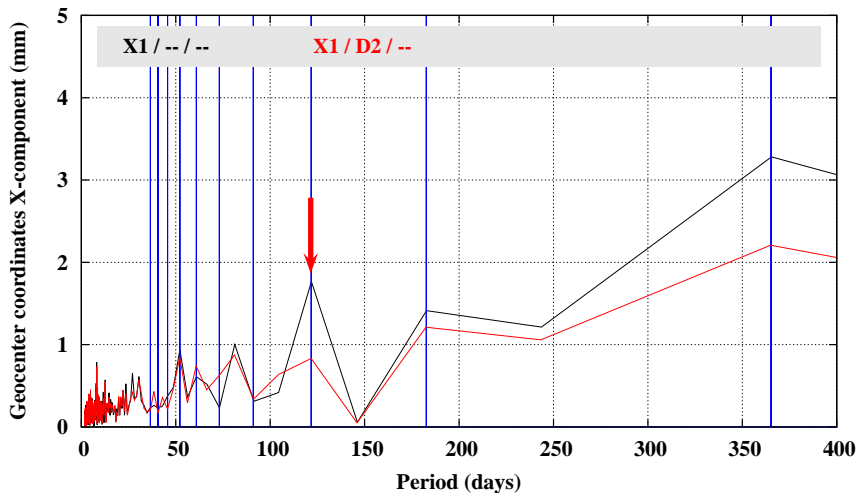
Impact on the Geocenter Estimates

Spectra from geocenter estimates: X component



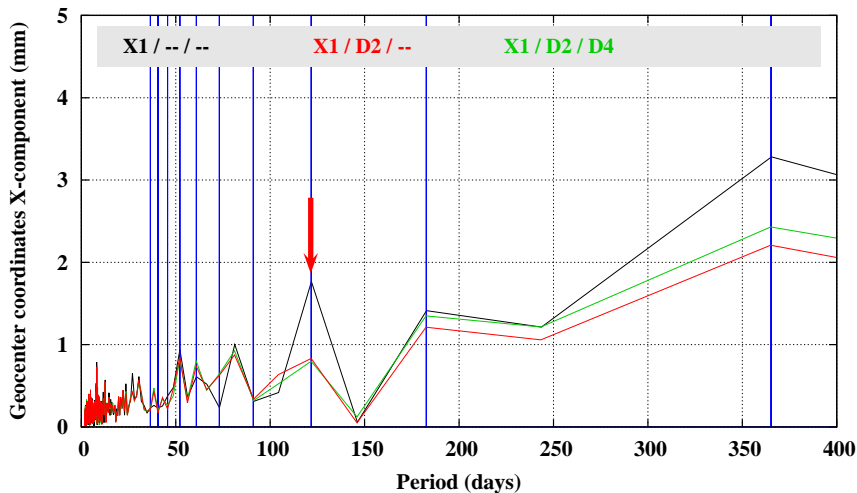
Impact on the Geocenter Estimates

Spectra from geocenter estimates: X component



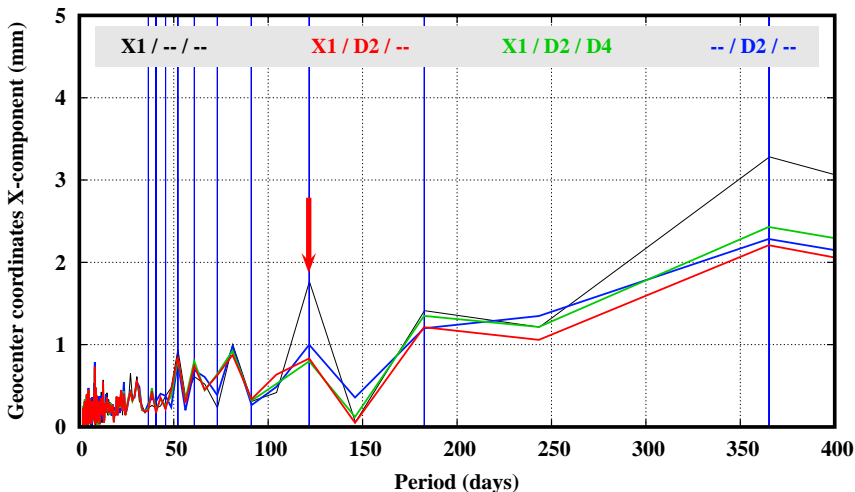
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Spectra from geocenter estimates: X component



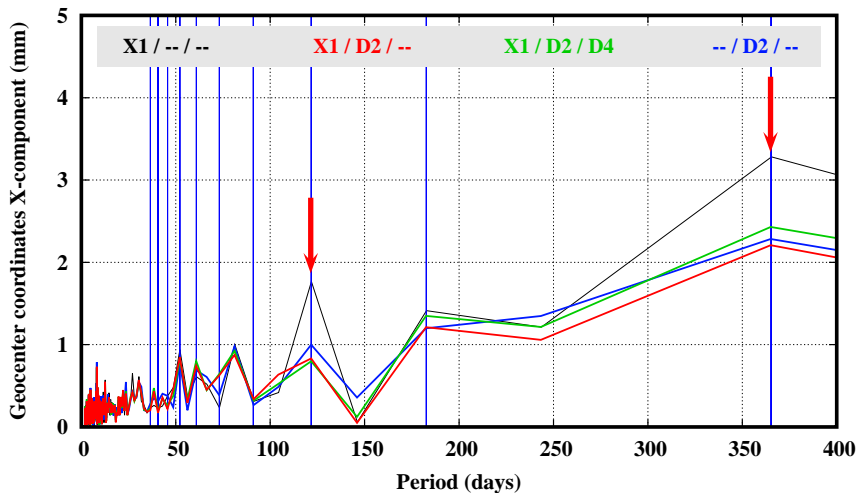
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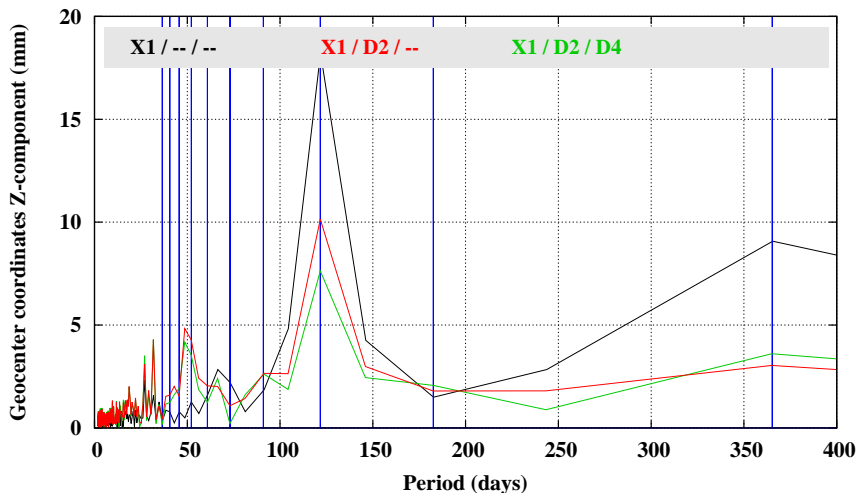
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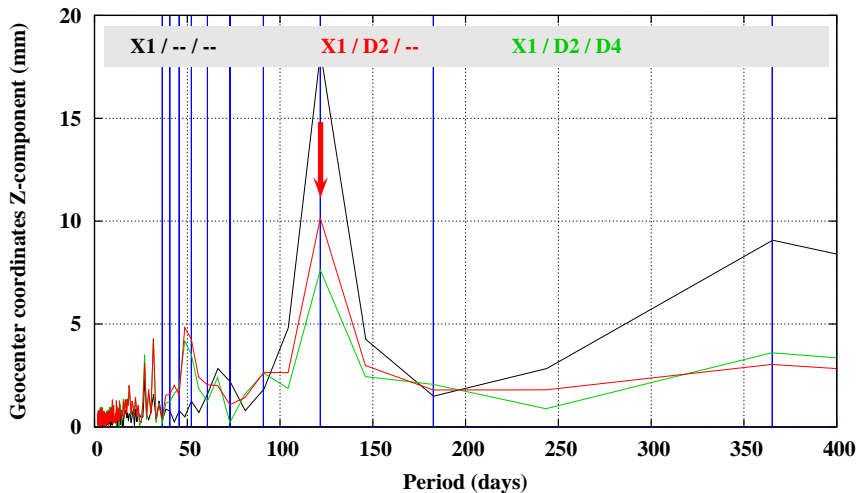
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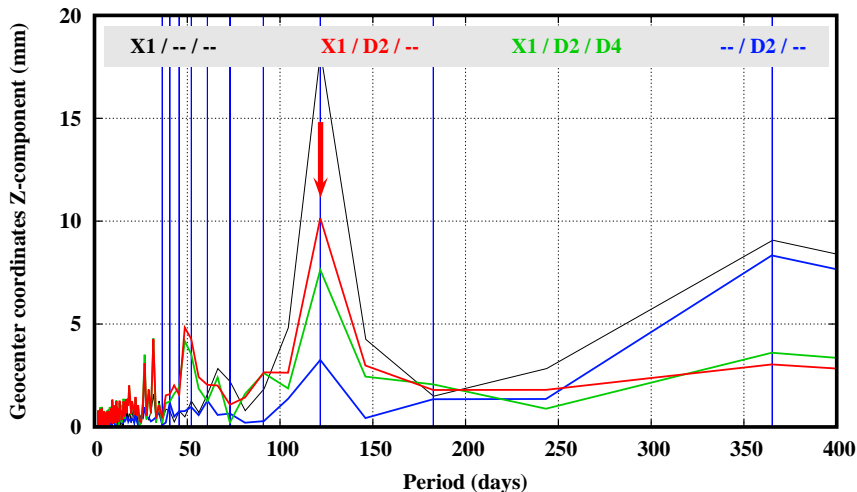
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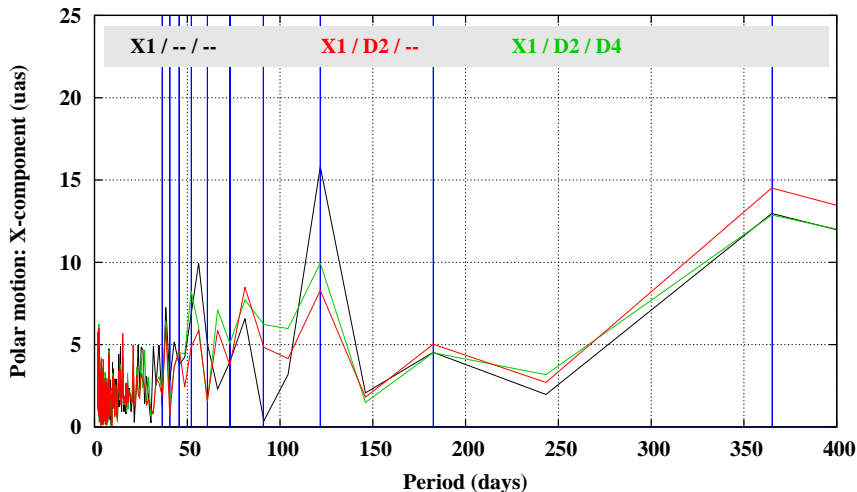
Impact on the Geocenter Estimates

Spectra from geocenter estimates: Z component



Impact on the Earth Rotation Parameters

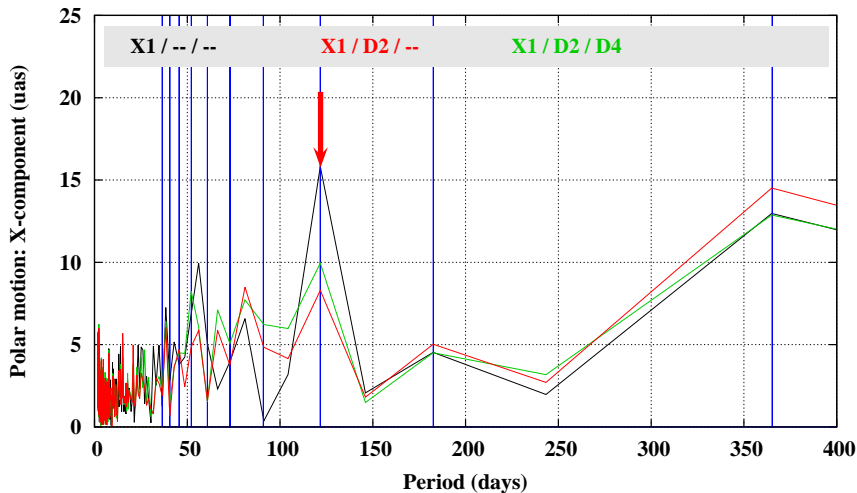
Spectra from ERP solution: Polar motion – X



Differences w.r.t. IERS C04 series (related to ITRF2008) has been analysed.

Impact on the Earth Rotation Parameters

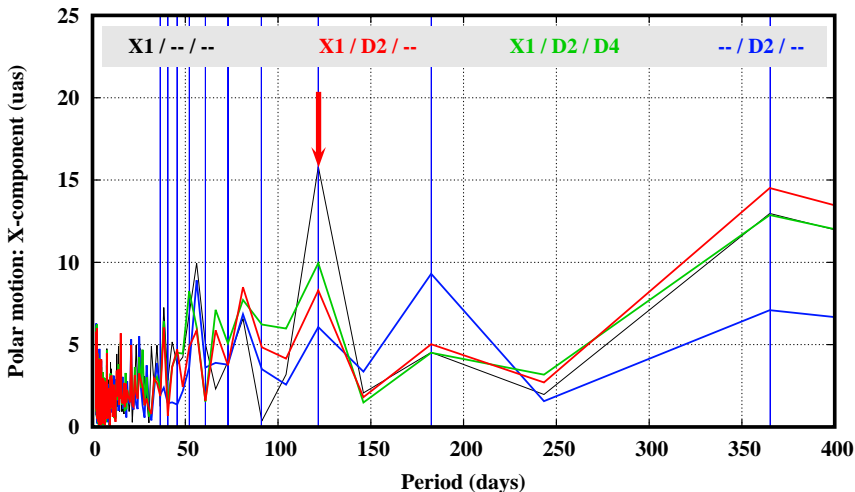
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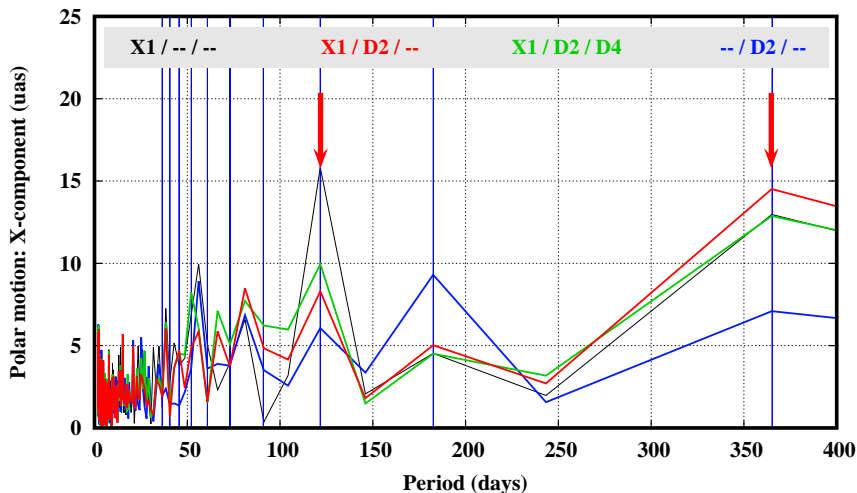
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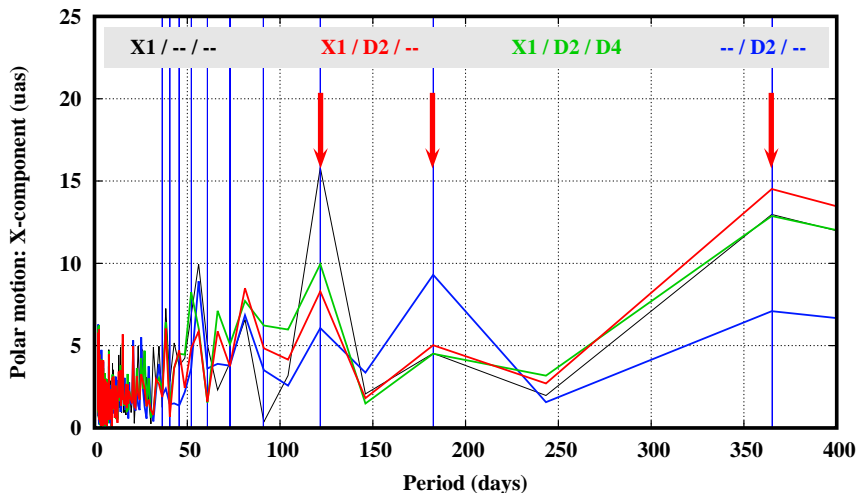
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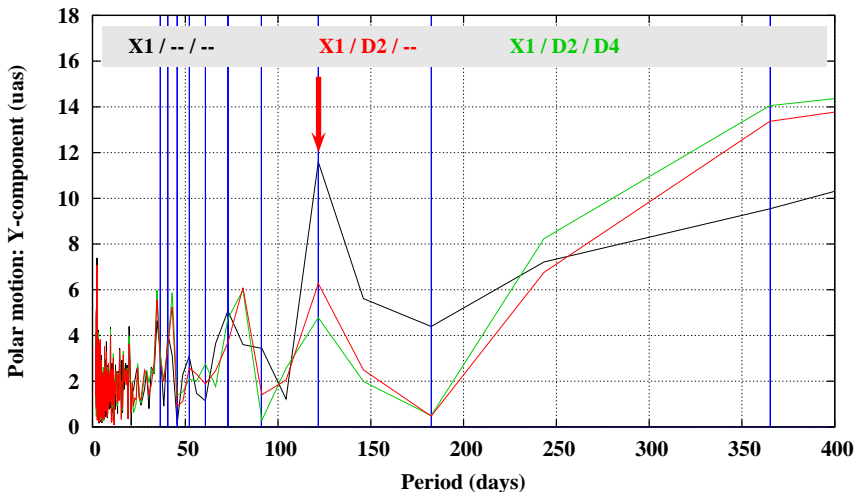
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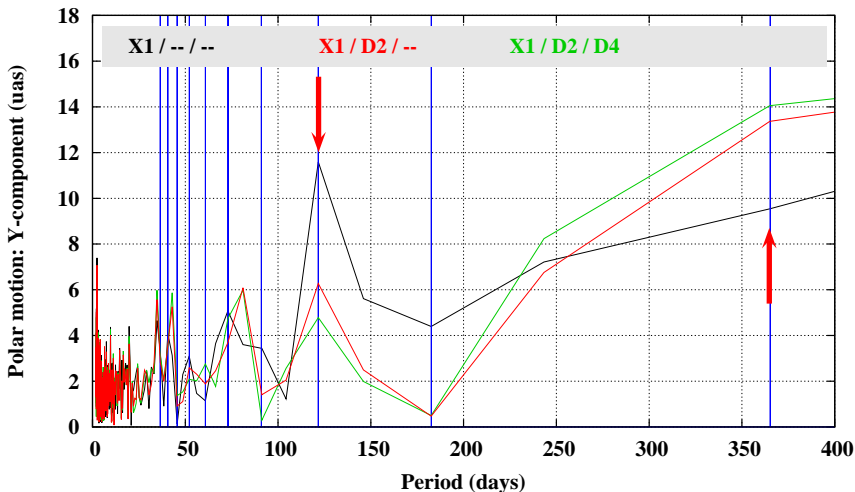
Spectra from ERP solution: Polar motion – Y



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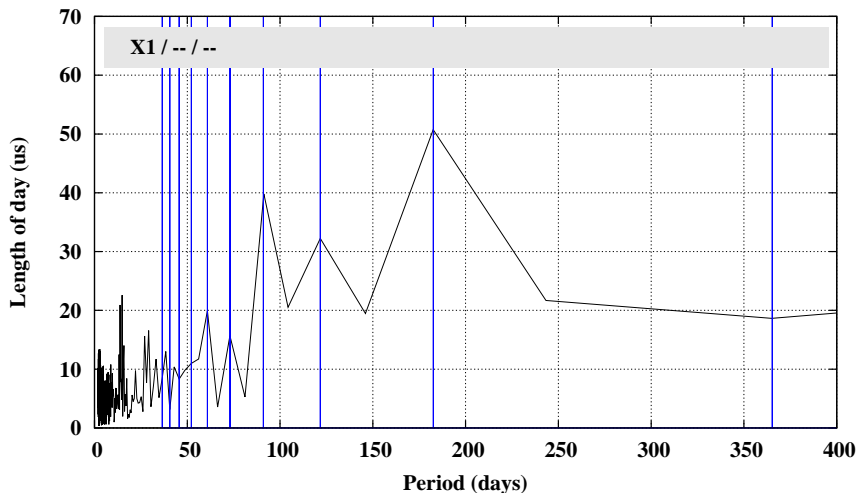
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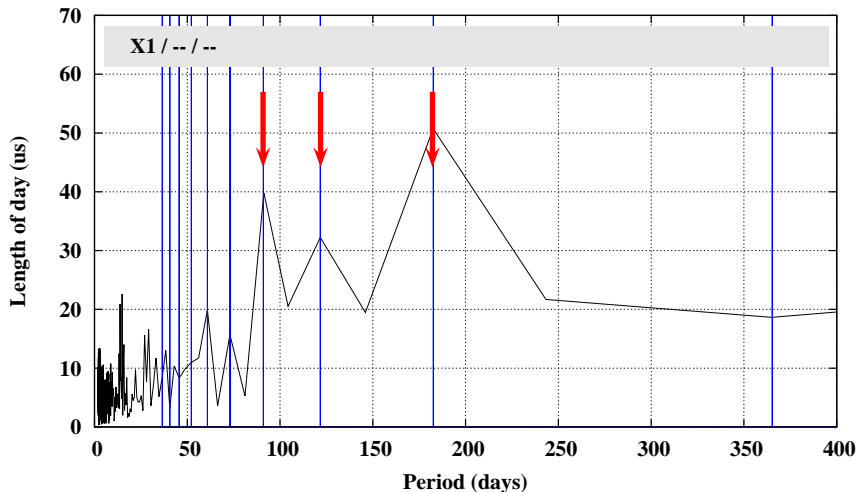
Spectra from ERP solution: length of day



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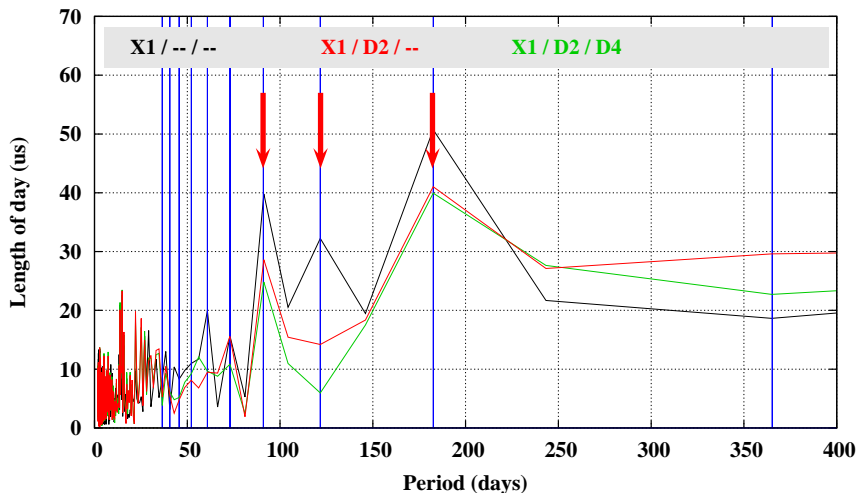
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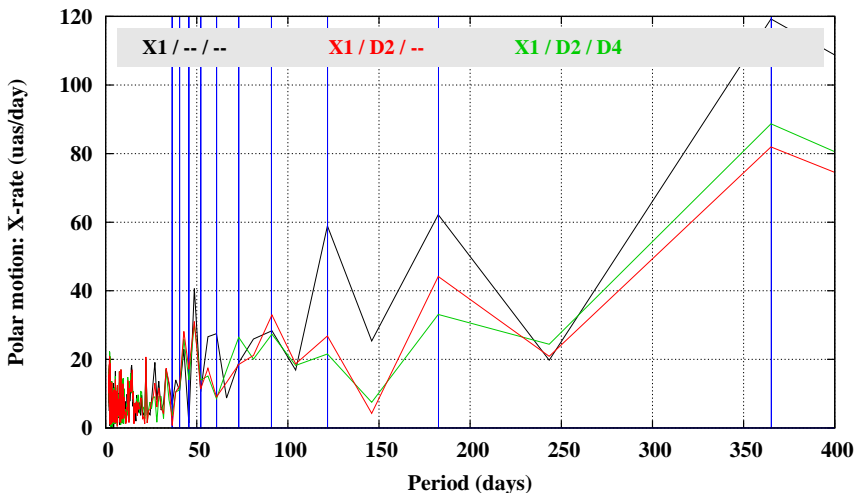
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Impact on the Earth Rotation Parameters

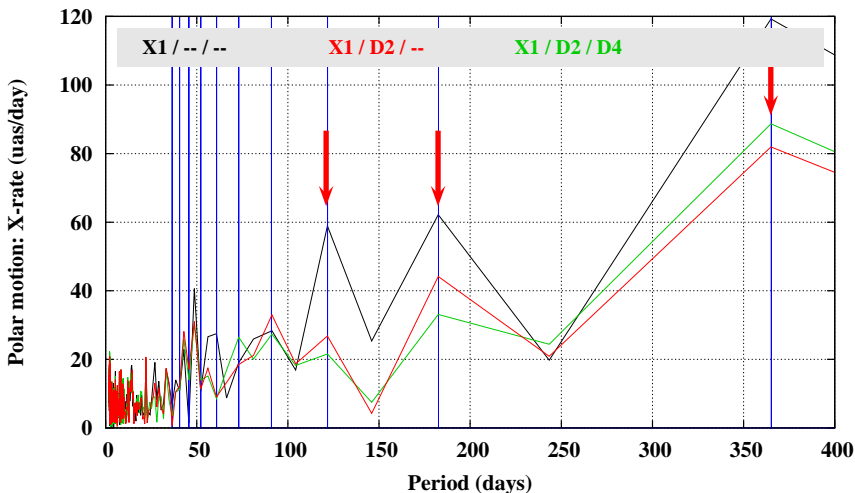
Spectra from ERP solution: Polar motion – X rate



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Impact on the Earth Rotation Parameters

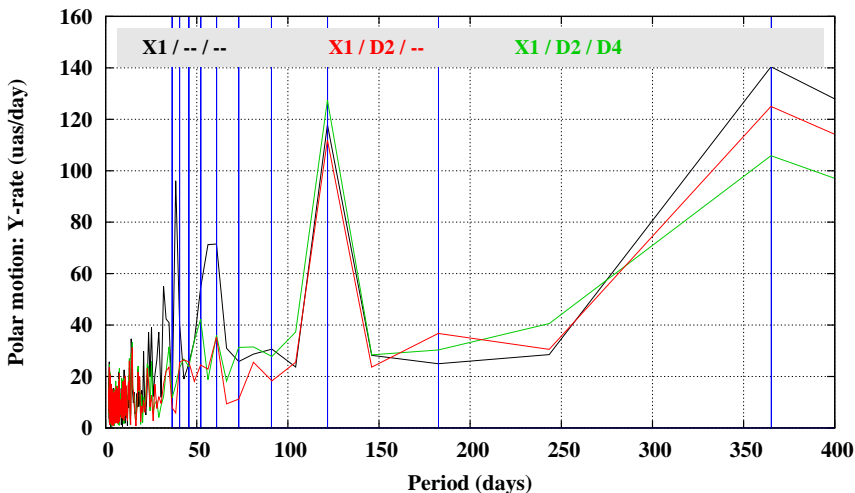
Spectra from ERP solution: Polar motion – X rate



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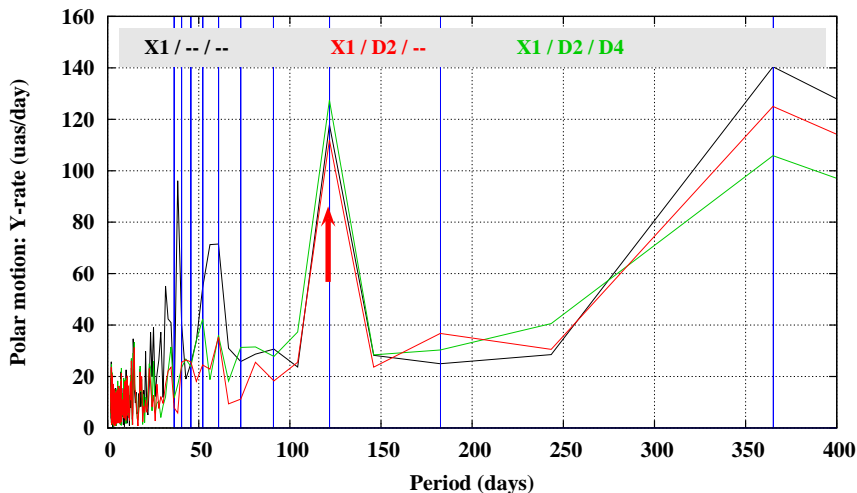
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- The analysis of the Earth rotation parameters shows a similar effect (apart from \dot{Y}).
- The most promising orbit parameter setup is: X_1 , D_2 (and D_4).

Long-Arc Solutions

Strategy for the long-arc solution

Classical approach to generate three-day solutions at CODE:

NEQ from day -1

NEQ from day ± 0

NEQ from day $+1$

ORB	ORB	ORB
ERP	ERP	ERP
CRD	CRD	CRD
TRP	TRP	TRP
⋮	⋮	⋮

Long-Arc Solutions

Strategy for the long-arc solution

Classical approach to generate three-day solutions at CODE:

NEQ for long-arc solution, day ± 0



ORB

ERP

CRD

TRP

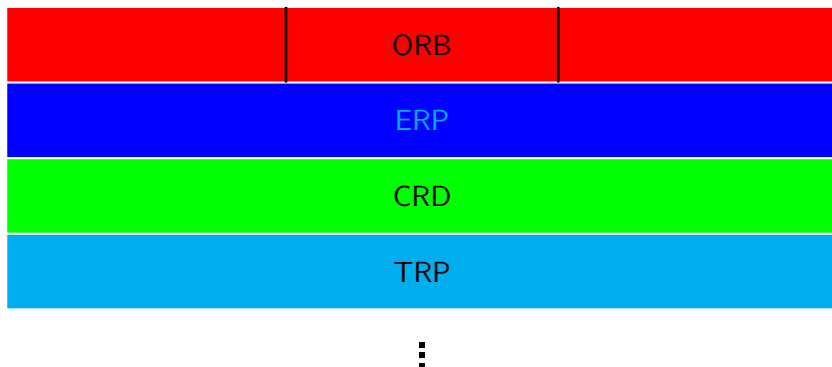
⋮

Long-Arc Solutions

Strategy for the long-arc solution

Classical approach to generate three-day solutions at CODE:

NEQ for long-arc solution, day ± 0



Long-Arc Solutions

Strategy for the long-arc solution

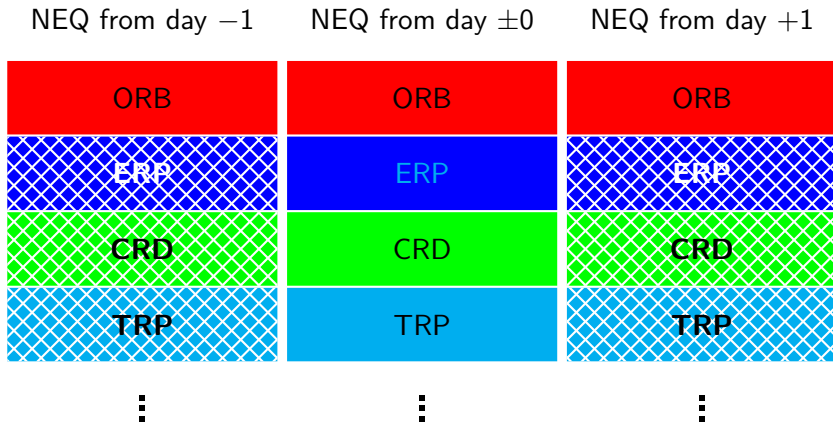
Alternative approach for a three-day long-arc solutions:

NEQ from day -1	NEQ from day ± 0	NEQ from day $+1$
ORB	ORB	ORB
ERP	ERP	ERP
CRD	CRD	CRD
TRP	TRP	TRP
⋮	⋮	⋮

Long-Arc Solutions

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Alternative approach for a three-day long-arc solutions:

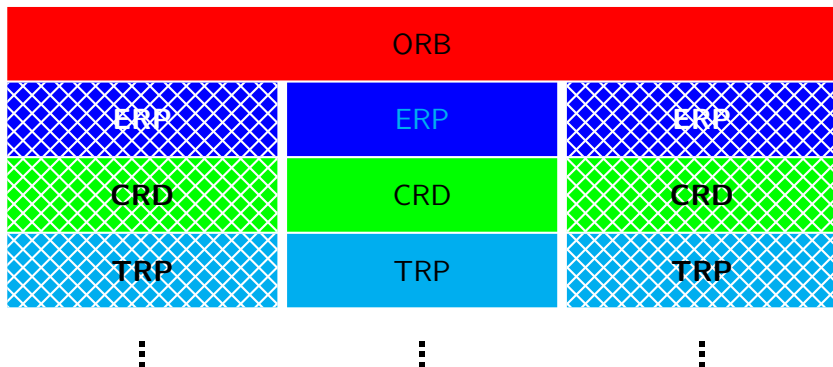


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NEQ for long-arc solution, day ± 0

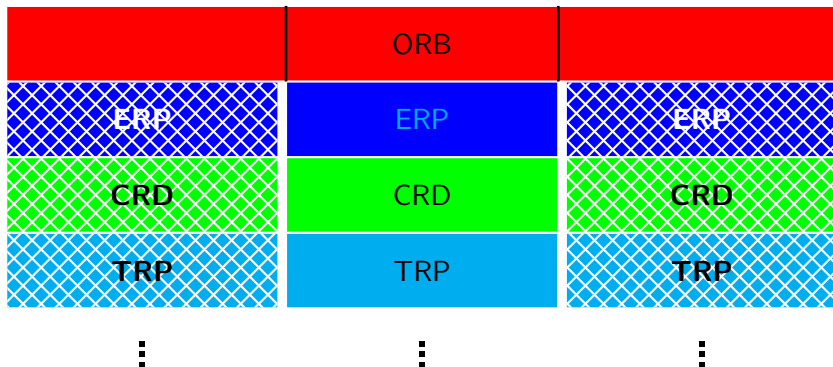


Long-Arc Solutions

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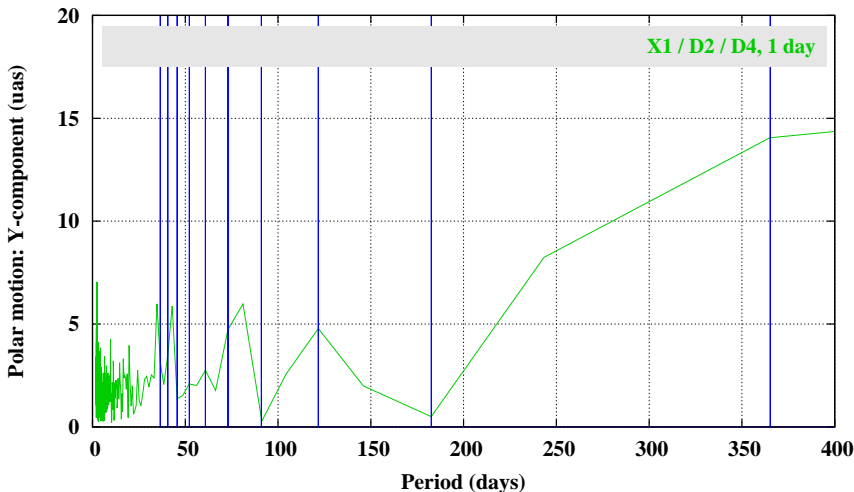
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Impact on the Earth Rotation Parameters

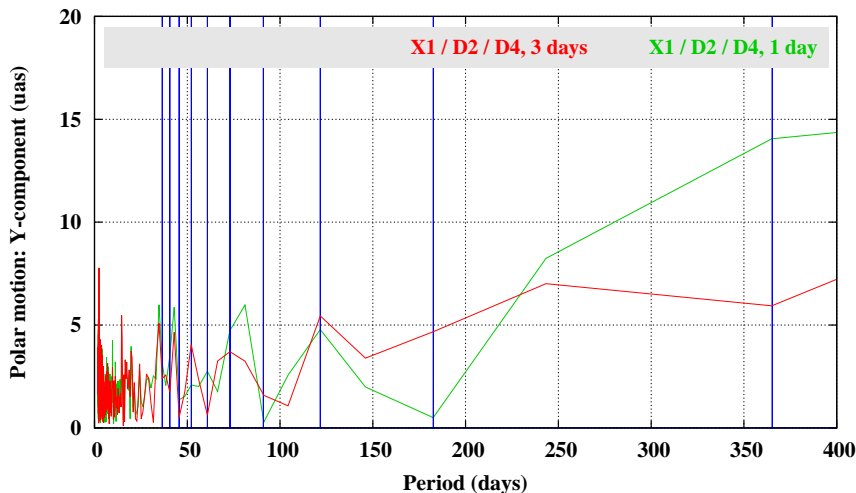
Spectra from ERP solution: Polar motion – Y



Differences w.r.t. IERS C04 series (related to ITRF2008) has been analysed.

Impact on the Earth Rotation Parameters

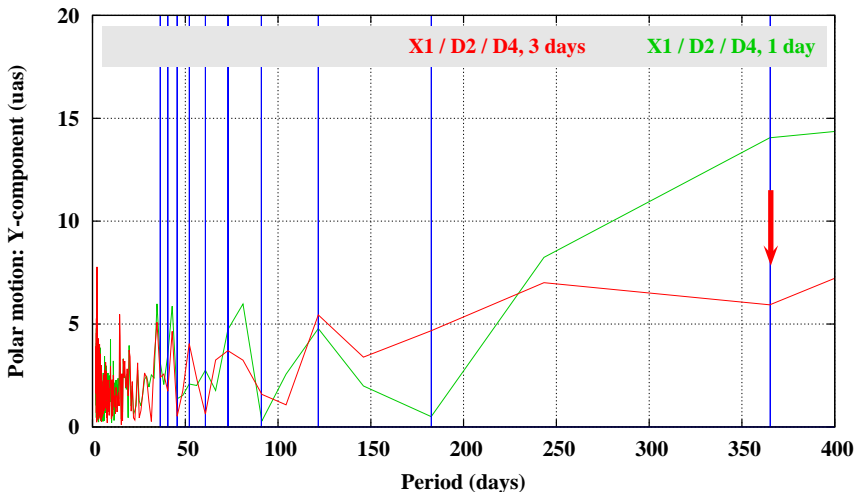
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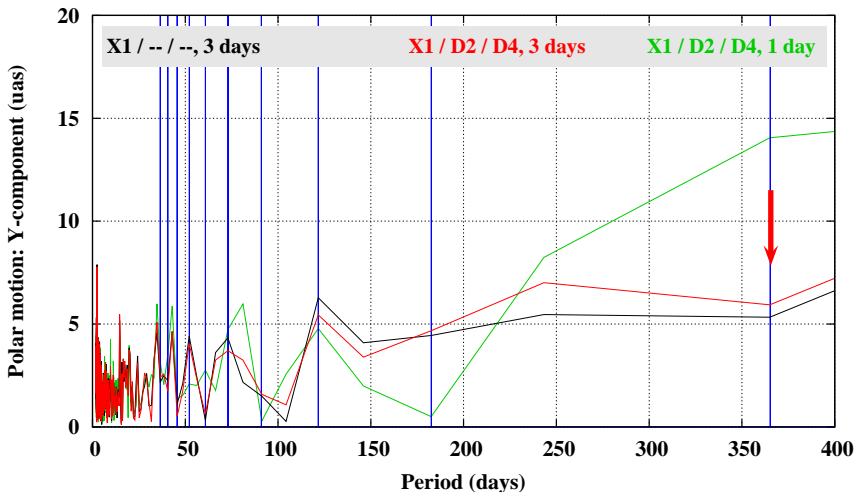
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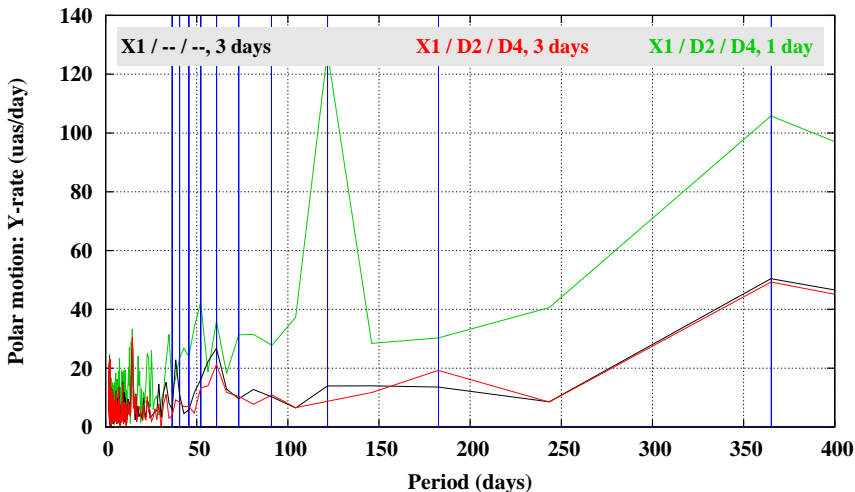
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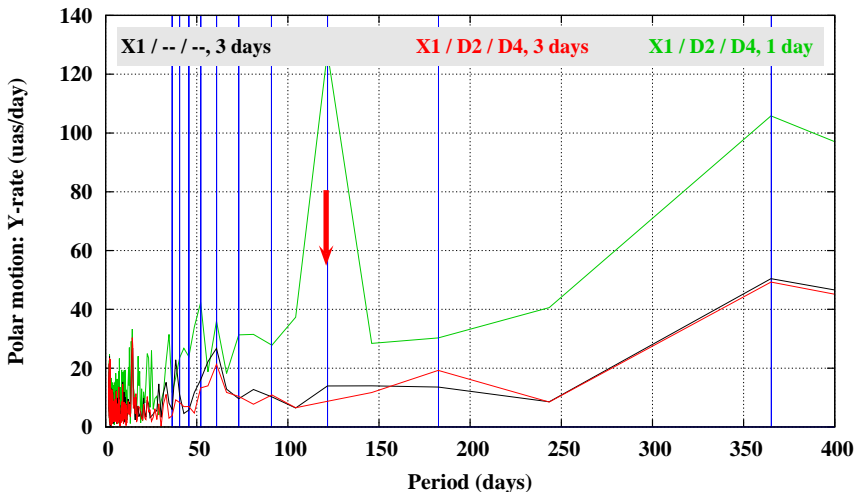
Spectra from ERP solution: Polar motion – Y rate



Differences w.r.t. IERS C04 series (related to ITRF2008) has been analysed.

Impact on the Earth Rotation Parameters

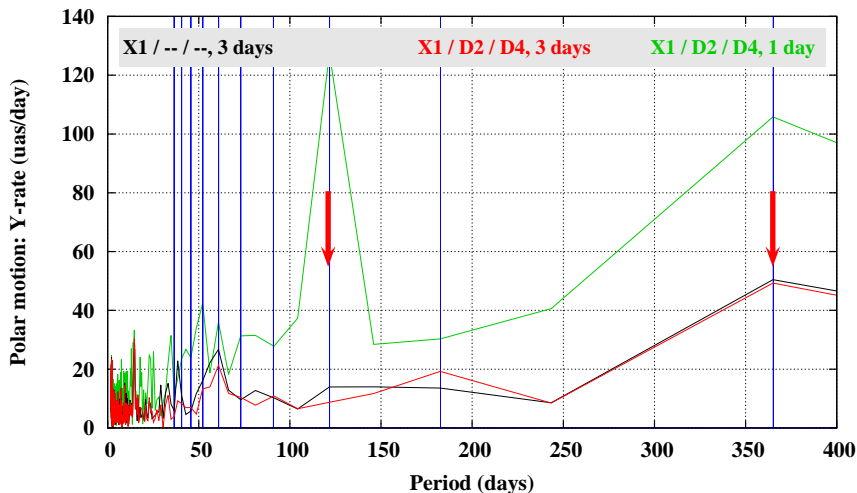
Spectra from ERP solution: Polar motion – Y rate



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Impact on the Earth Rotation Parameters

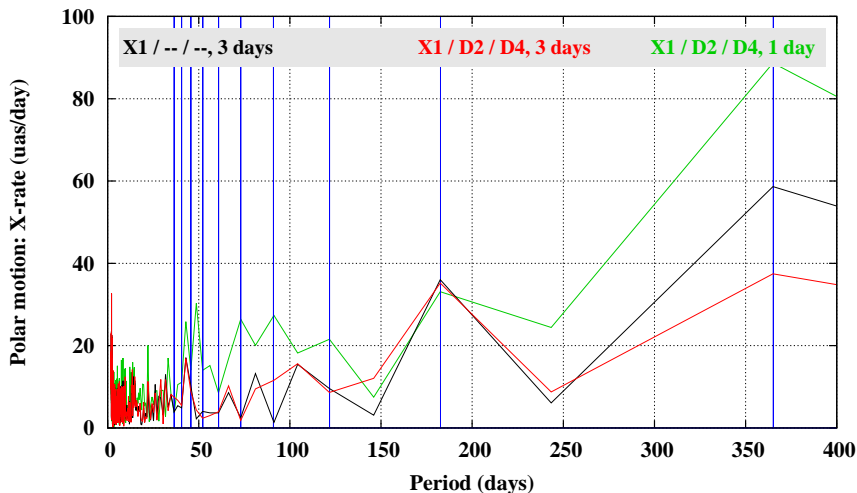
Spectra from ERP solution: Polar motion – Y rate



Differences w.r.t. IERS C04 series (related to ITRF2008) has been analysed.

Impact on the Earth Rotation Parameters

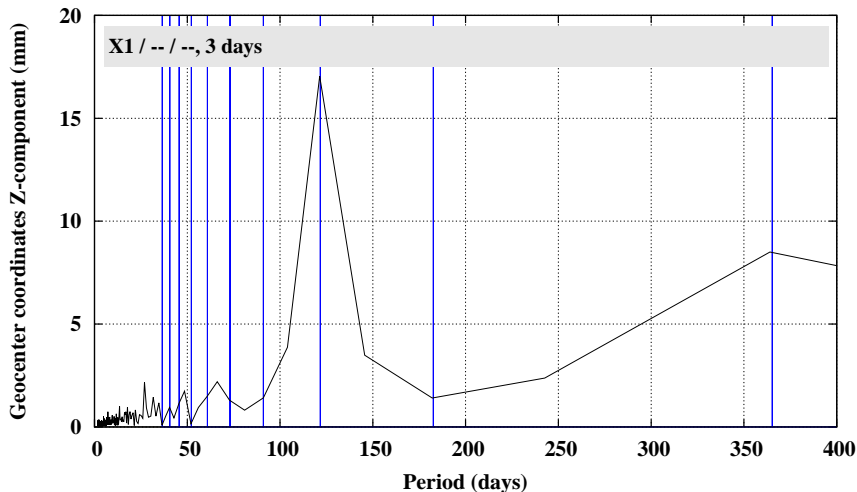
Spectra from ERP solution: Polar motion – X rate



Differences w.r.t. IERS C04 series (related to ITRF2008) has been analysed.

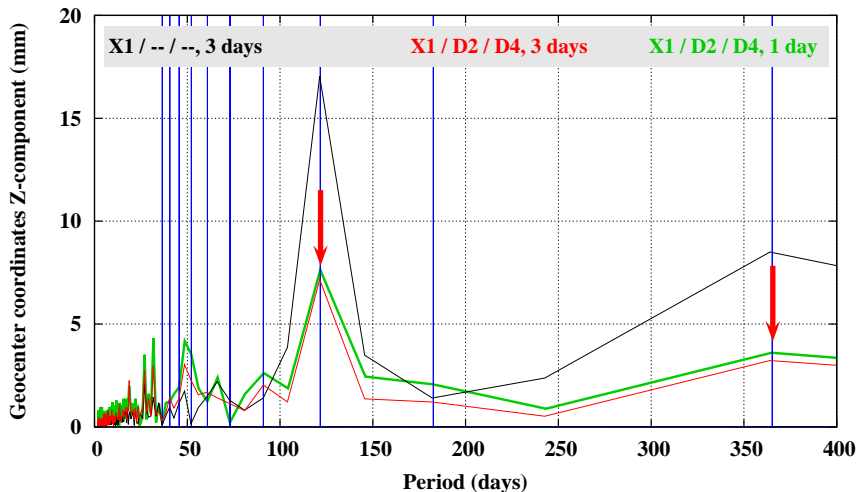
Impact on the Geocenter Estimates

Spectra from geocenter estimates: Z component



Impact on the Geocenter Estimates

Spectra from geocenter estimates: Z component



Conclusions

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- The improvement for the rates in polar motion (\dot{X} and \dot{Y}) and in the length of day component are remarkable.
- In the geocenter series the orbit parametrization is more important than the arc length.
- **CODE is currently preparing the transfer of this new orbit model into its operational IGS processing.**

THANK YOU for your attention



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